

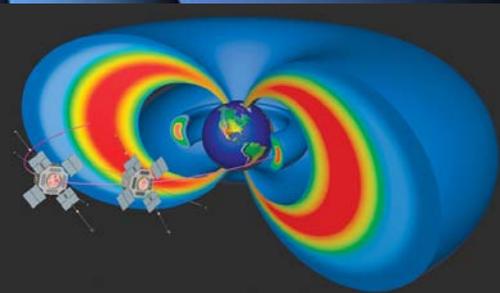
Recent Findings on Characteristics of Chorus waves observed on THEMIS

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We acknowledge all THEMIS
team for providing data.

UCLA

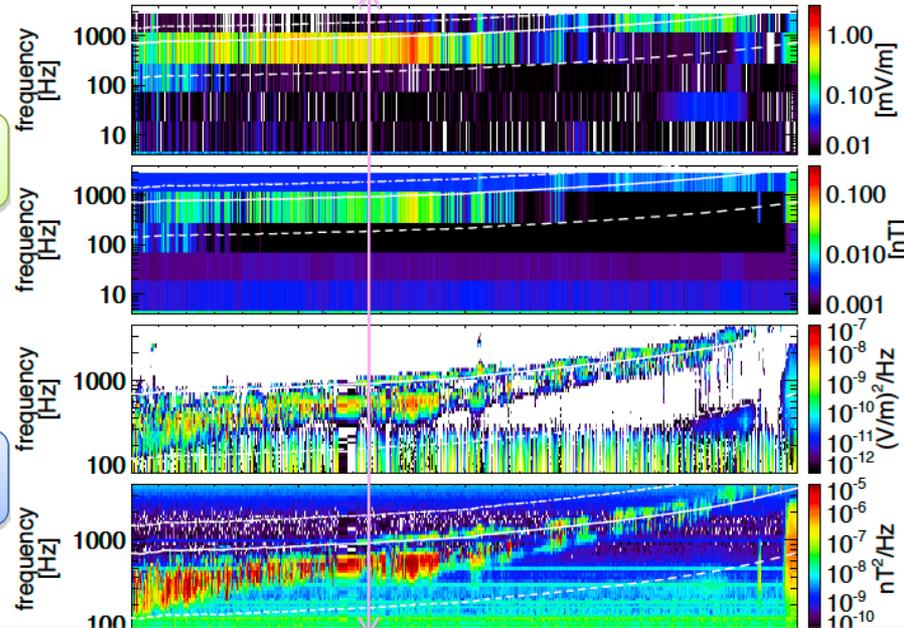
Outline

- ✧ **Global distribution of chorus wave amplitudes, in particular the large amplitude chorus**
- ✧ **Wave normal angle distribution of lower-band and upper-band chorus**
- ✧ **Evolution of chorus amplitude distribution at various phases of CIR-driven storms**
- ✧ **Typical properties of various whistler-mode waves**

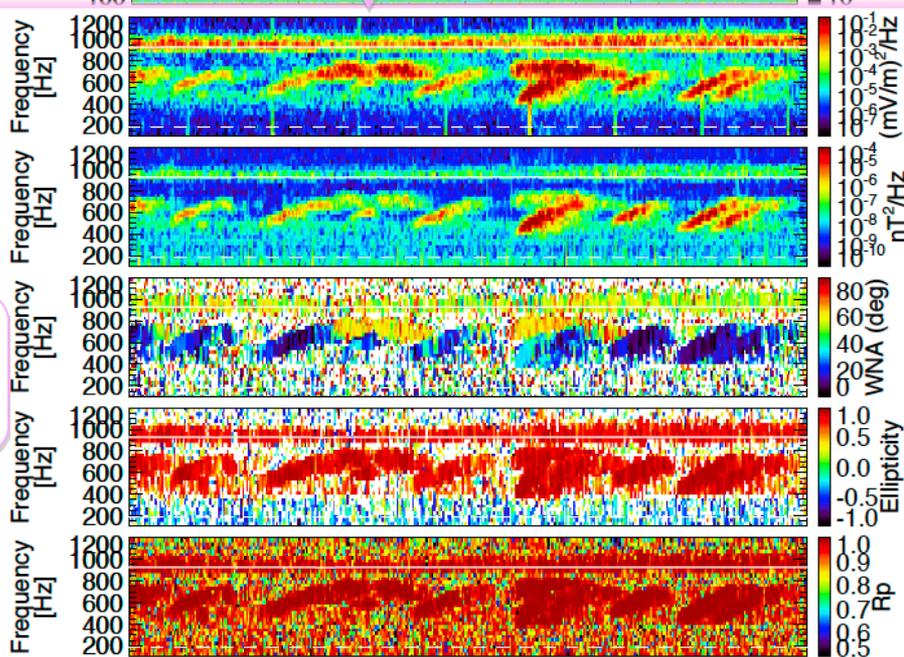
THEMIS Wave Data

- THEMIS A, D, E; outside pp and inside mp; 5-10 R_E ; all MLT; mostly $|MLAT| < 20^\circ$
- FBK data (2007/07/01-present) 4 sec time resolution
- FFF data (2010/05/01-present) high-resolution wave spectra data
- Waveform data (2007/07/01-present) sampling rate: up to ~16 kHz

FBK



FFF

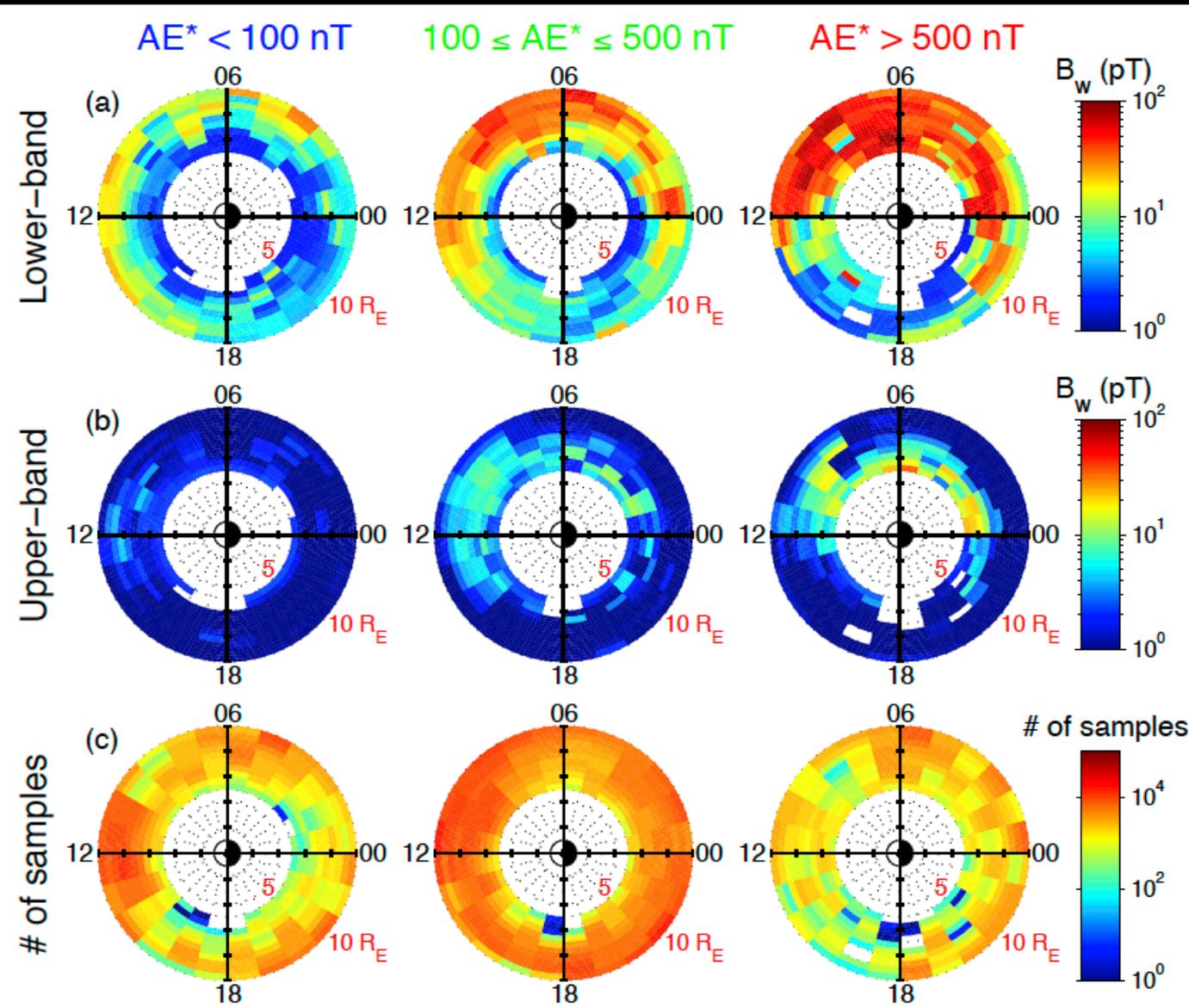


wave form

LAT	-1.2	-1.2	-1.2
L	8.3	8.3	8.3
MLT	9.3	9.3	9.3
Seconds	04	06	08

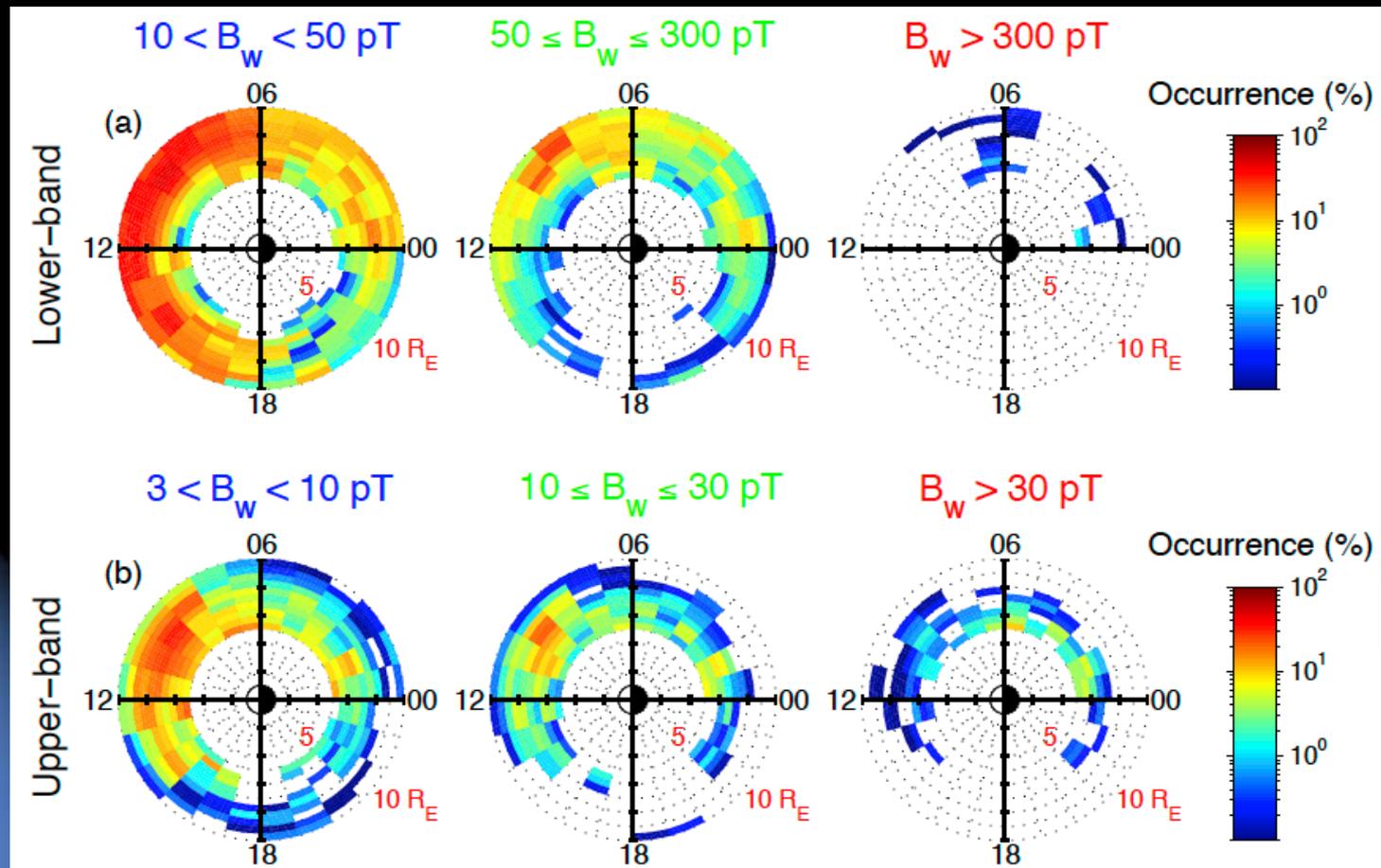
2011 Jan 10 2125:

Global distribution of chorus B_w (FFF)



- B_w depends on AE^*
- B_w near midnight is more closely related to AE^* .
- On the dayside, modest LB is observed even during quiet times.
- UB is much weaker and more confined to lower L (<8).
- Both LB and UB are observed from premidnight to afternoon

Occurrence of chorus waves for various wave amplitudes (FFF)

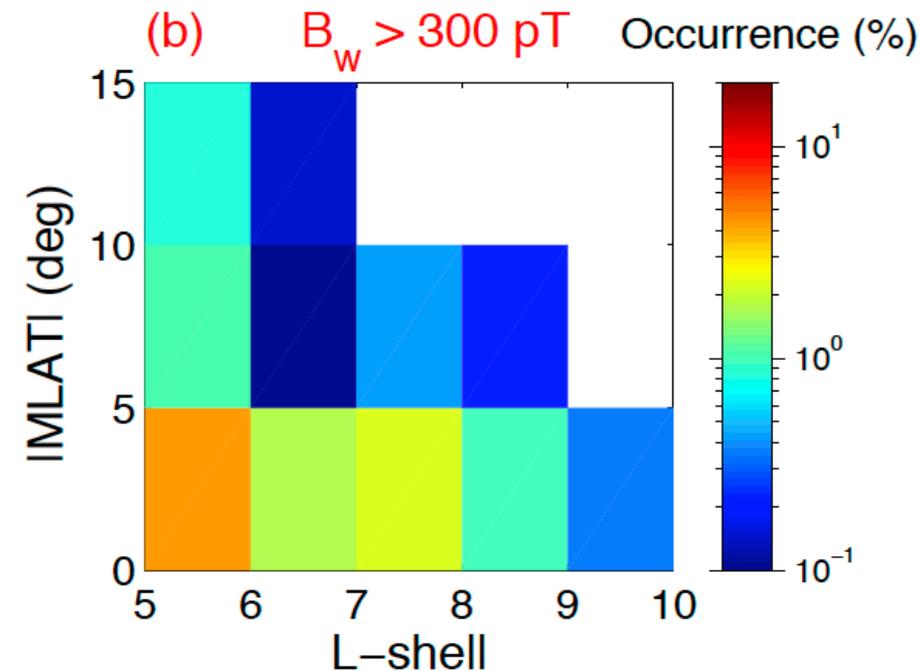
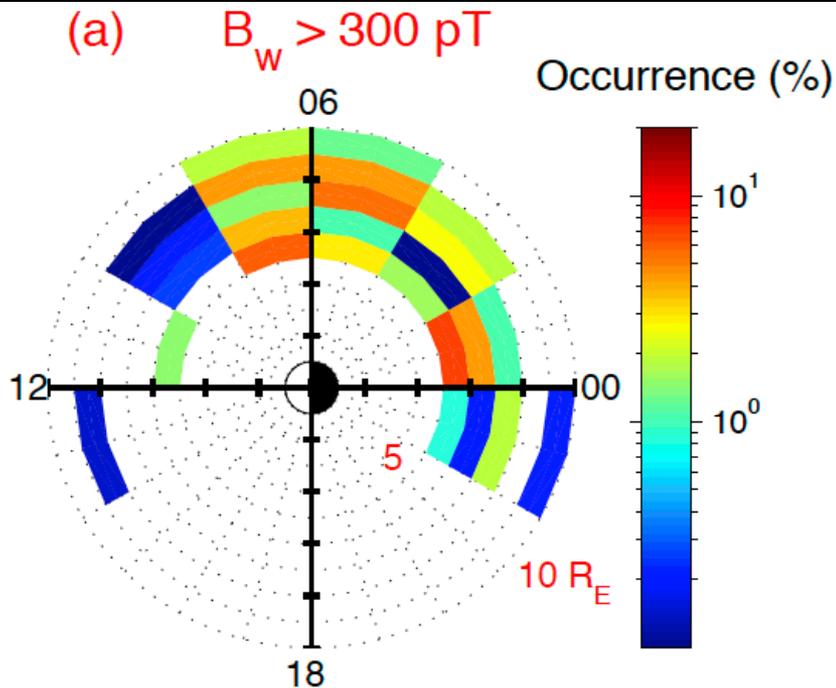


- **LB:** occurrence rate of large amplitude chorus (> 300 pT) is very small, while modest chorus is distributed over a broad MLT range with higher occurrence.
- **UB:** strong chorus is from pre-midnight to dawn, while weak chorus extends further into afternoon.

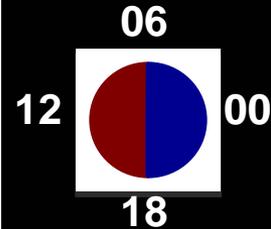
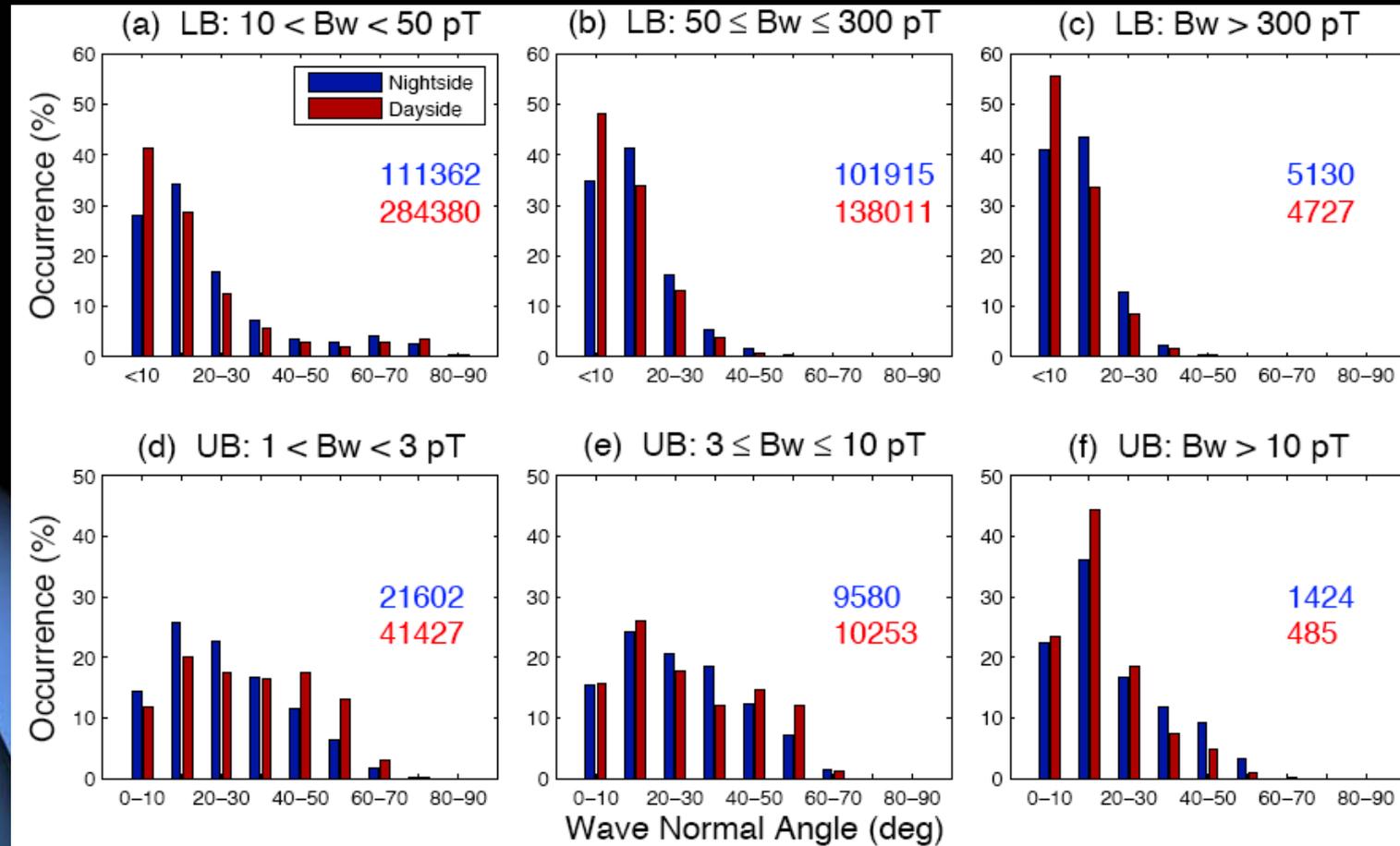
[Li et al., 2011]

Occurrence of large-amplitude chorus (waveform)

- Occurrence is high from premidnight to postdawn up to a few %.
- Preferentially observed at lower L-shells
- Higher occurrence closer to the equator.
- Occurrence rate is higher than that obtained from FFF data, which may confirm that high resolution waveform data are essential to capture large amplitude chorus.



Wave normal angle distribution for various wave amplitude ranges (waveform)



LB

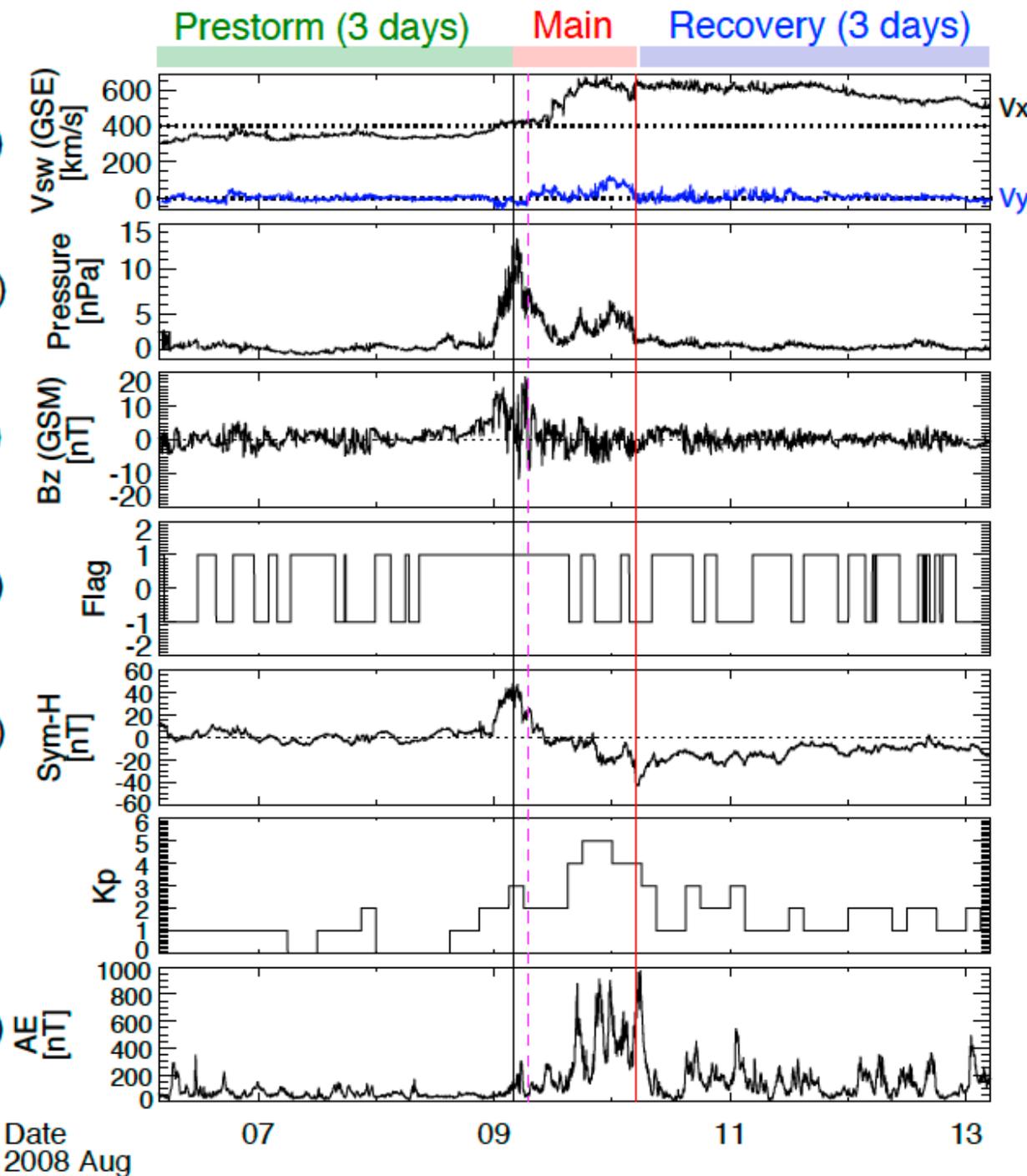
UB

- ✧ **LB:** Majority of WNAs are less than 20° ; WNA tends to be smaller as B_w increases; WNA on the dayside is generally smaller than on the nightside.
- ✧ **UB:** WNA of UB is generally larger than LB, distributed in a broad range, while strong UB still peaks at $10-20^\circ$.

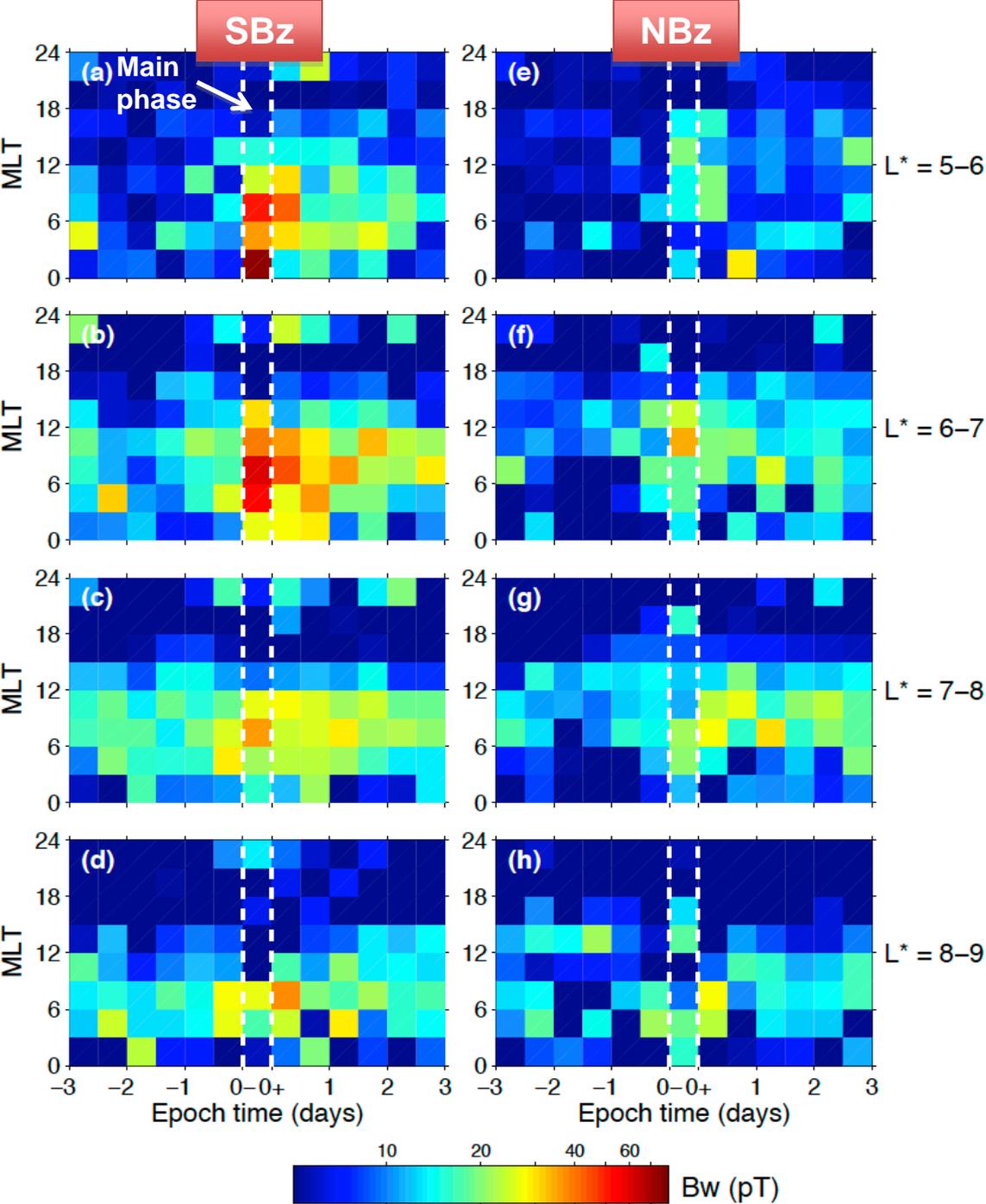
One example of CIR storm

Superposed epoch analysis

- ✧ 3 days before + main phase + 3 days after
- ✧ 2007/07/01 – 2011/10/01
- ✧ 72 CIR storms
- ✧ Categorize CIR events by IMF B_z southward or northward



Flag
1: mean(B_z) in the previous 2.5 hour > 0
-1: < 0



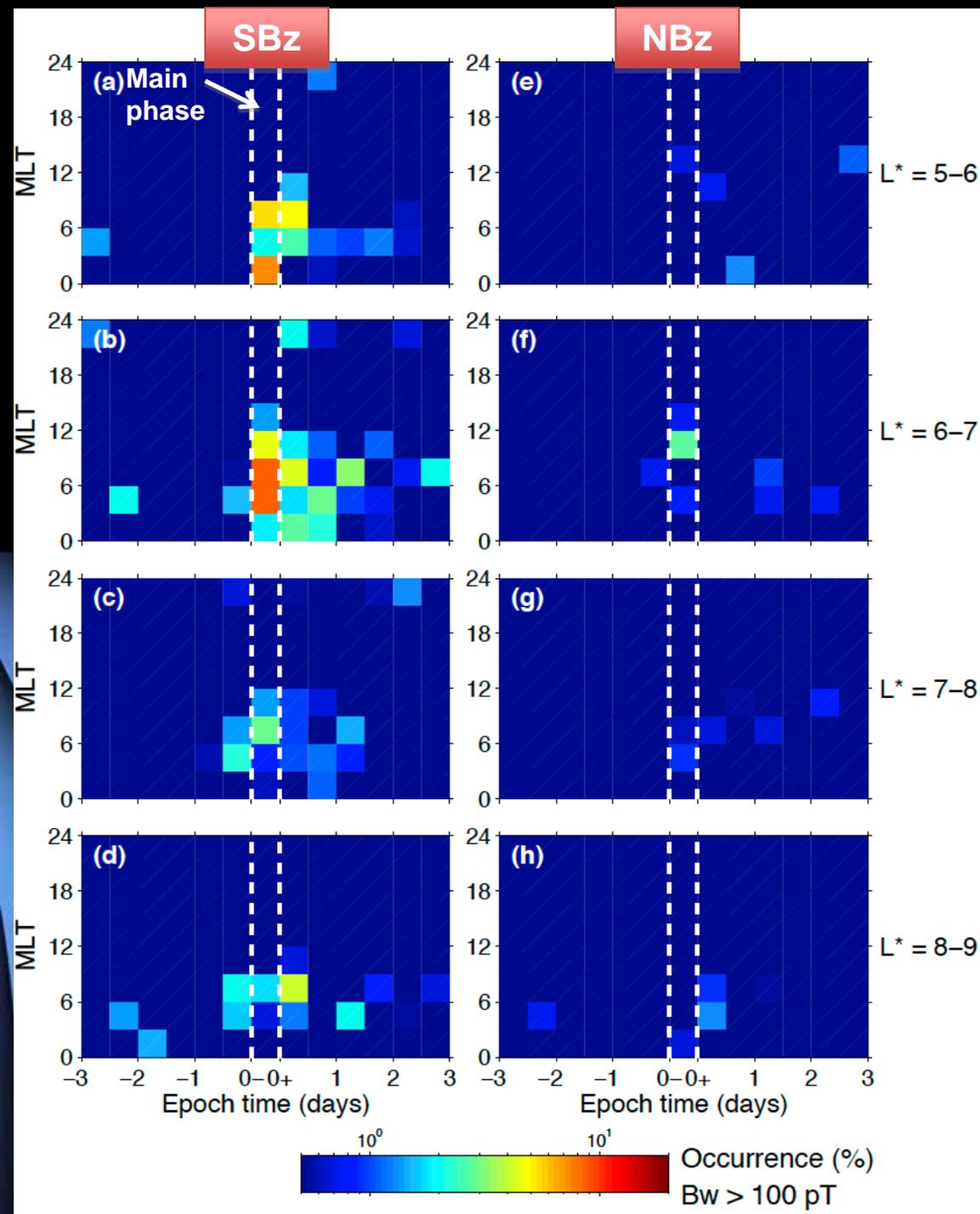
Evolution of chorus wave amplitude (B_w) during CIR-driven storms – superposed epoch analysis (FBK)

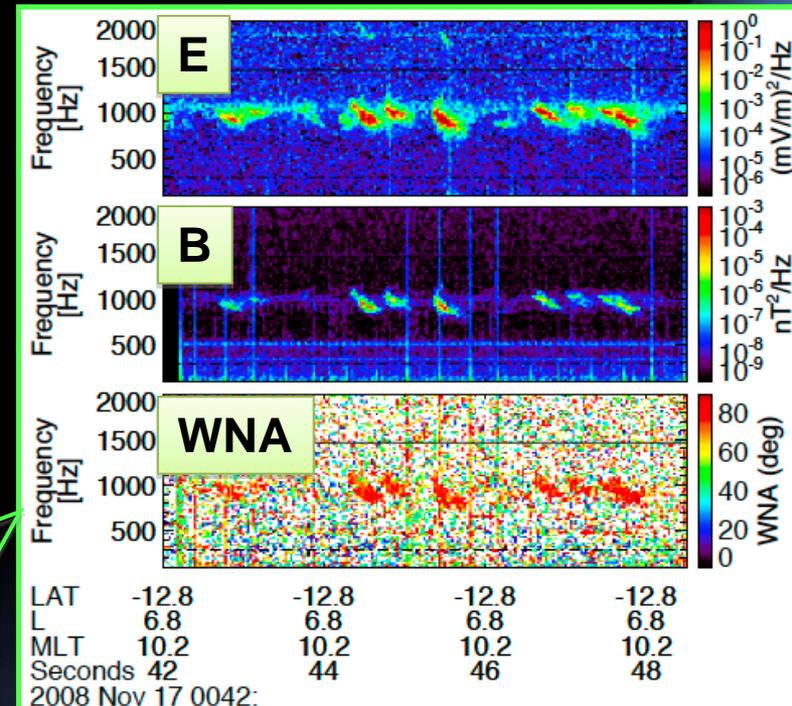
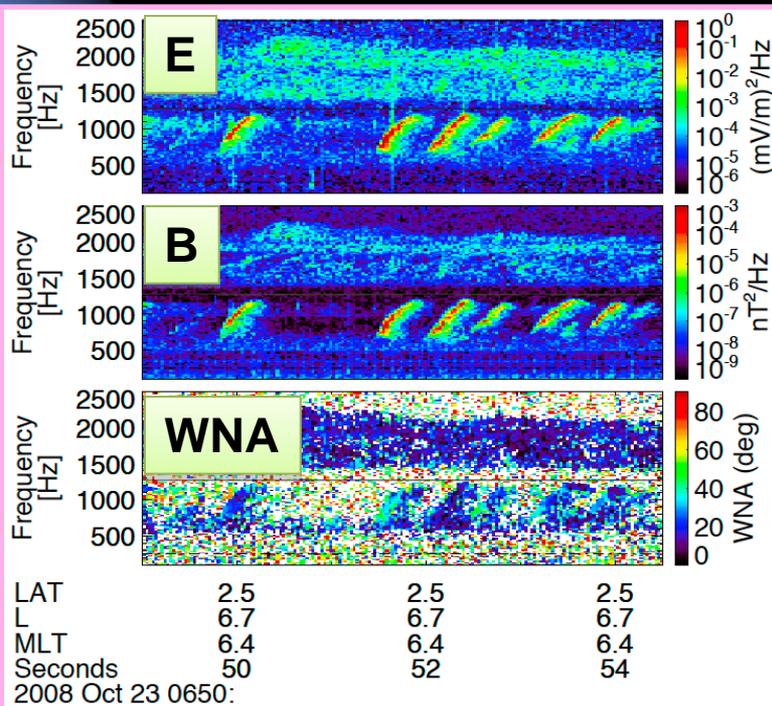
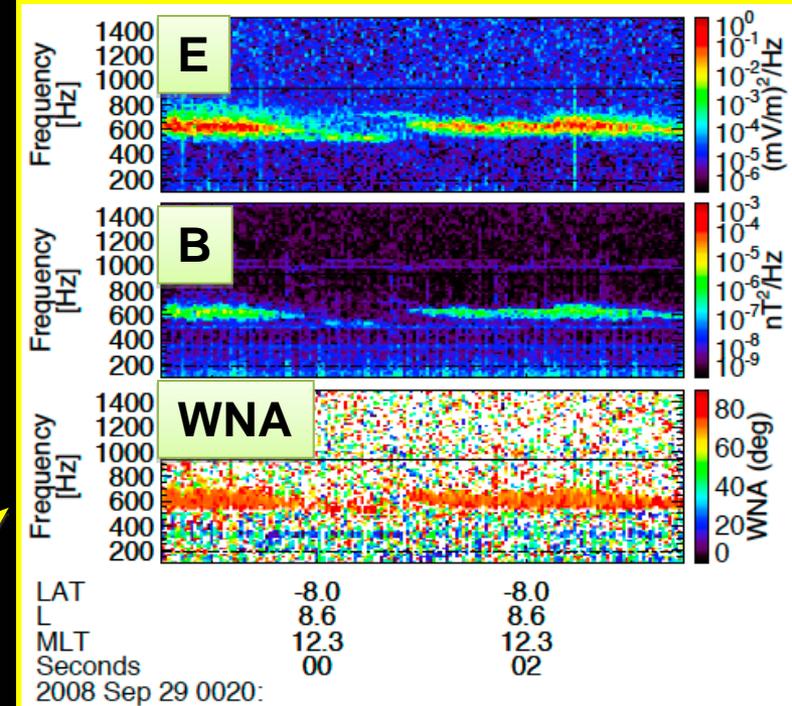
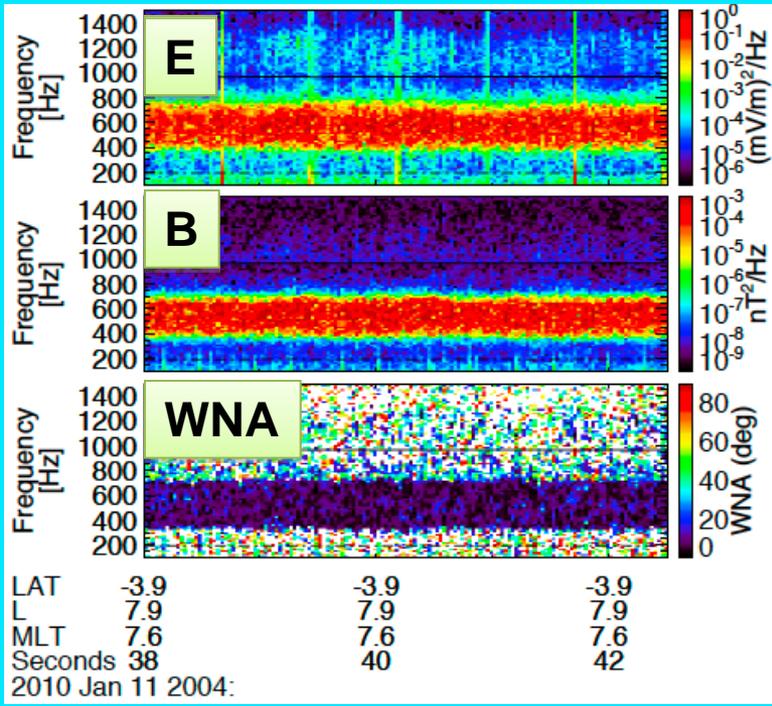
- ✧ In SBz, B_w is weak during prestorm, increases at 0-12 MLT during the main phase, and gradually decreases during the subsequent recovery phase.
- ✧ B_w during the recovery phase is stronger than that during prestorm.
- ✧ At $L^* < 7$, B_w at 0-12 MLT peaks in the main phase.
- ✧ Compared to SBz category, in NBz category B_w is much weaker at $L^* < 7$, but is comparable at $L^* > 7$.

Evolution of occurrence rate of strong chorus for $B_w > 100$ pT (FBK)

Strong chorus waves are observed with highest occurrence rate at

- ◇ the main phase
- ◇ SBz
- ◇ 0-9 MLT
- ◇ $L^* < 7$



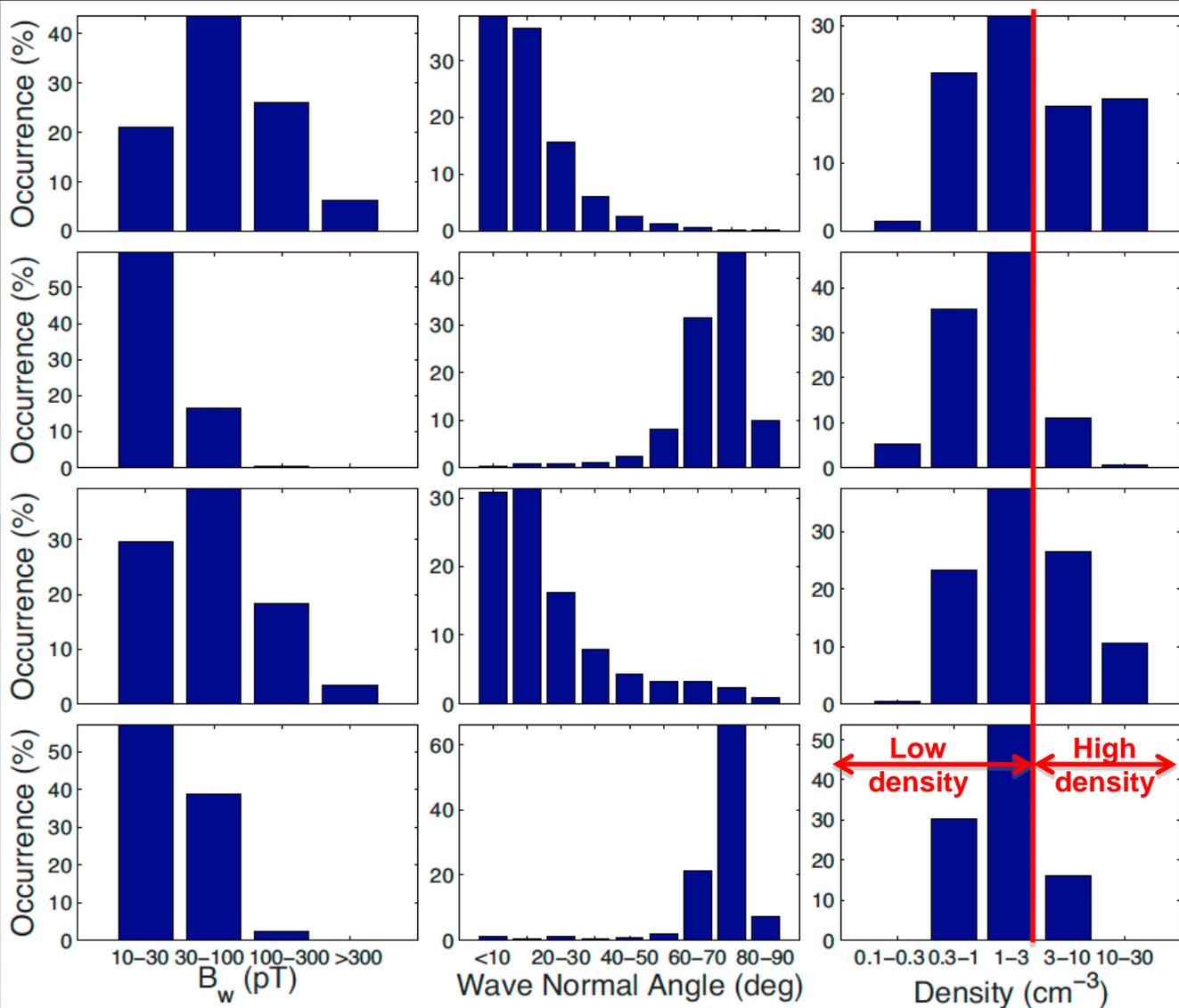


Typical properties of various emissions (**waveform**)

B_w

WNA

Density



Broad
band

Hissy-
like

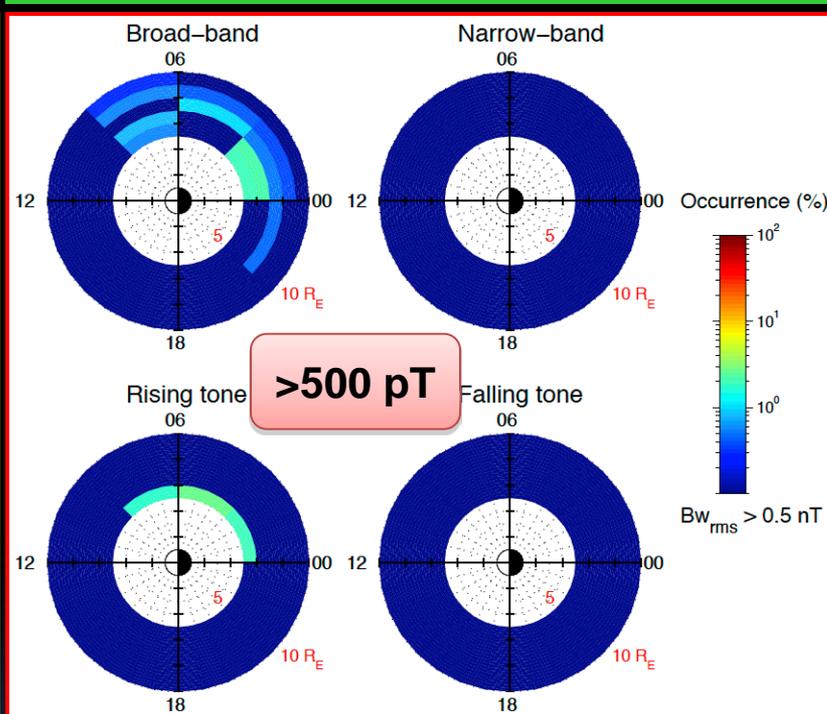
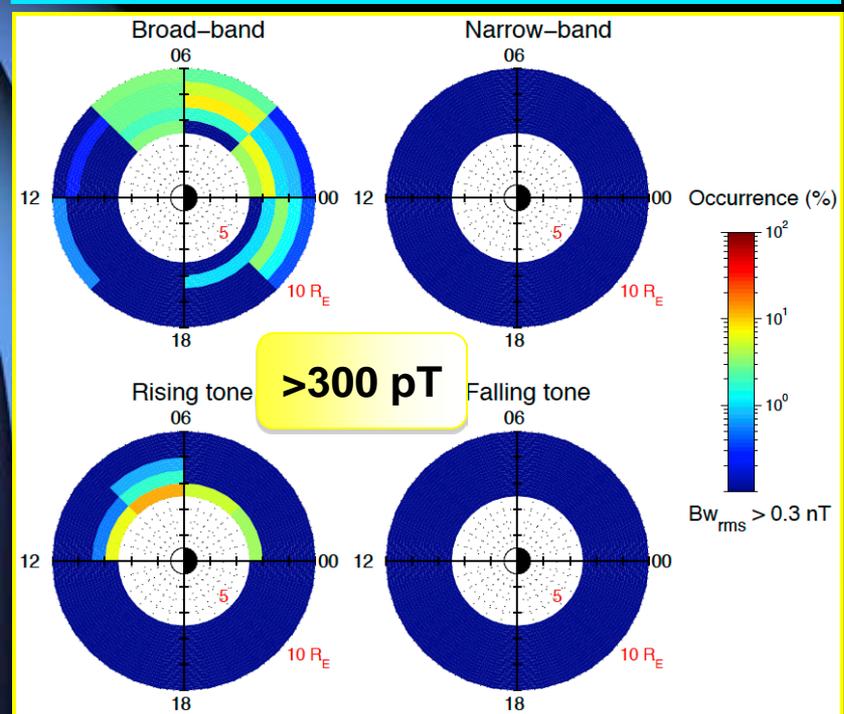
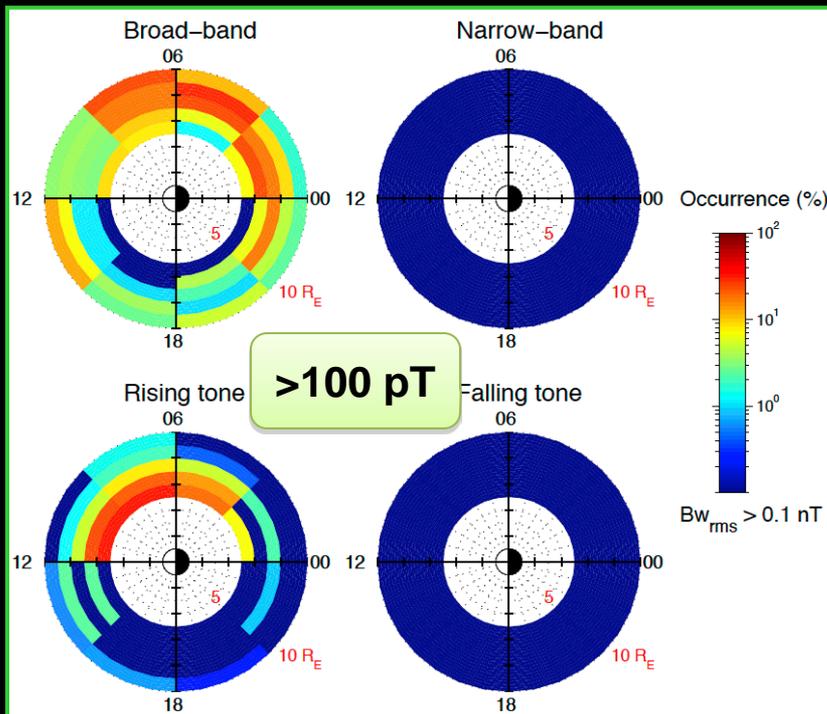
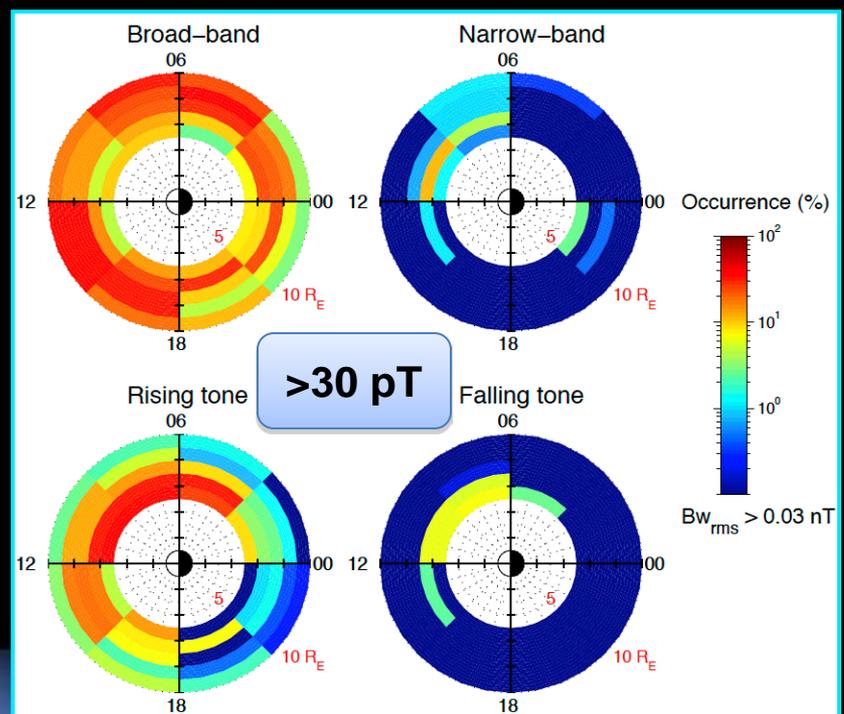
Narrow
band

Rising
tone

Discrete

Falling
tone

Occurrence rate of 4 types of emissions at various wave amplitudes in L-MLT domain (waveform)

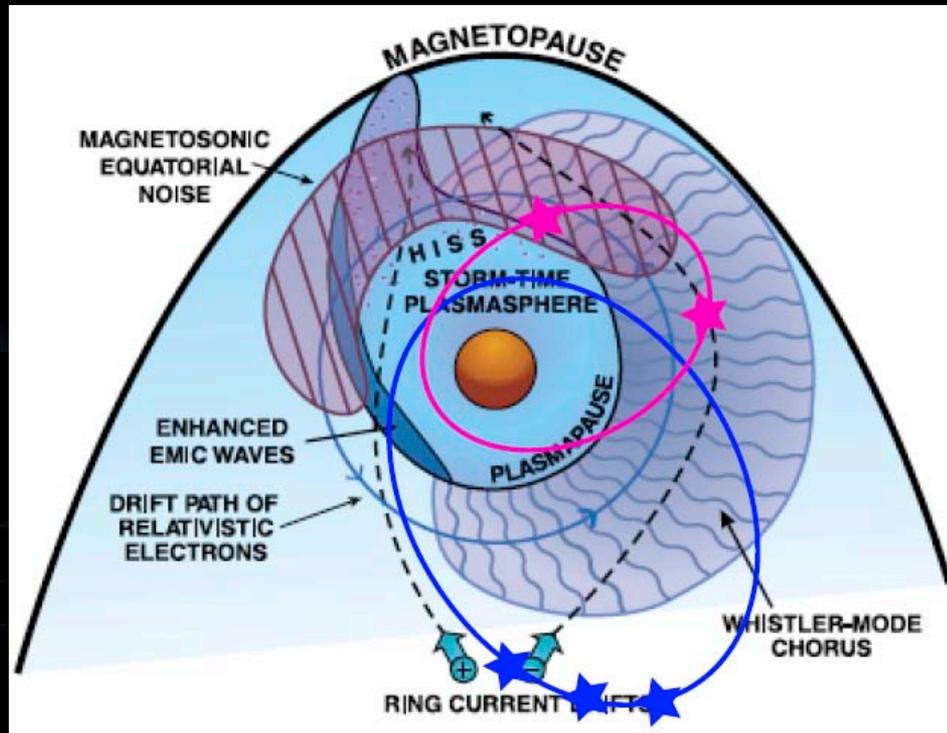


SUMMARY

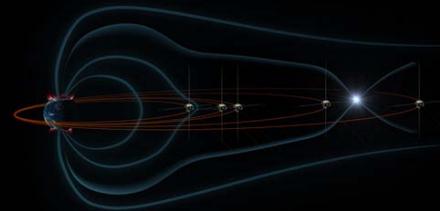
- ✧ Using THEMIS data, we provided more complete coverage of chorus wave amplitude distributions including large amplitude chorus, which has not been covered by previous satellites.
- ✧ Majority of WNAs of LB chorus are less than 20° and they tend to be smaller as B_w increases; WNA of UB is generally larger than LB, distributed in a broad range, while strong UB still peaks at $10\text{-}20^\circ$.
- ✧ During CIR-driven storms, chorus wave amplitudes peak in the main phase at $L^* < 7$ and are stronger under SBz than NBz IMF orientation.
- ✧ B_w of broad-band waves and rising tones are stronger than narrow-band and falling tones; broad-band and rising tones are typically quasi field-aligned ($<30^\circ$), while narrow-band and falling tones are preferentially very oblique. Broad-band waves are distributed in a broad range, while rising tones prefer to occur lower L.

OPEN QUESTIONS

- ✧ Generation of upper-band chorus and the gap at $0.5 f_{ce}$
- ✧ Generation mechanism of rising/falling tones and hissy-like emissions (what controls their emission types?)



THEMIS



RBSP

